

BLACK HOLES AND FUNDAMENTAL FIELDS, SCHOOL & WORKSHOP, LISBON, 1-5 JULY 2024

*Symmetry and nonlinearity
in black hole perturbation theory*

Lam Hui

Columbia University

Work done with: Joyce, Penco, Santoni, Solomon; Beren, Sun; Lagos; Mittman, Lagos, Stein+; Podo

Ladder structure from $SO(3,1)$ explains vanishing Love number

Quadratic QNMs are detectable and interesting



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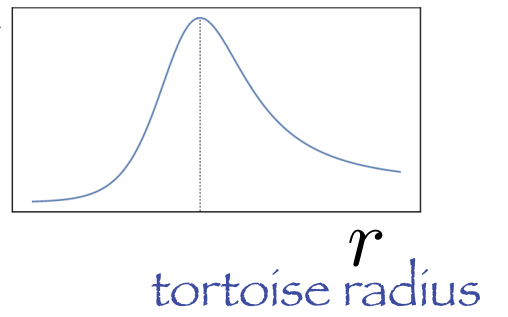
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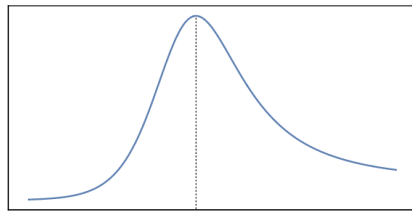
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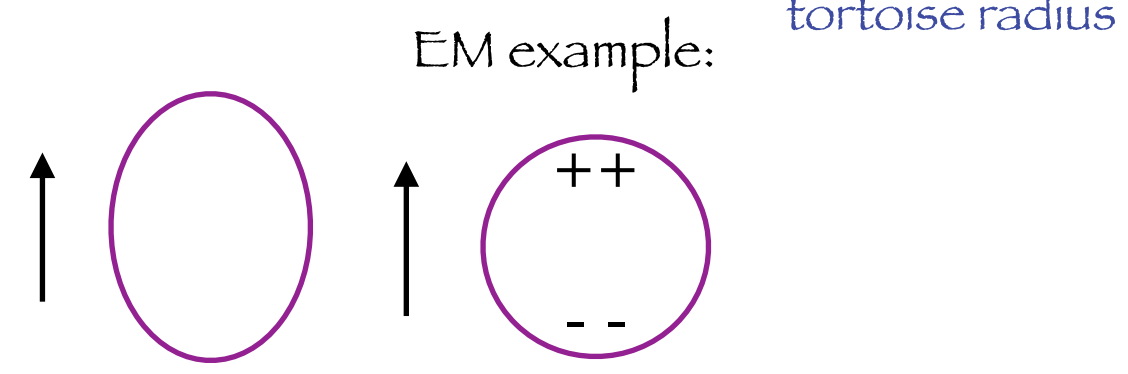
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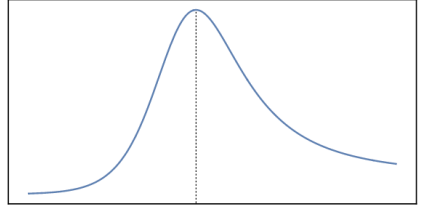


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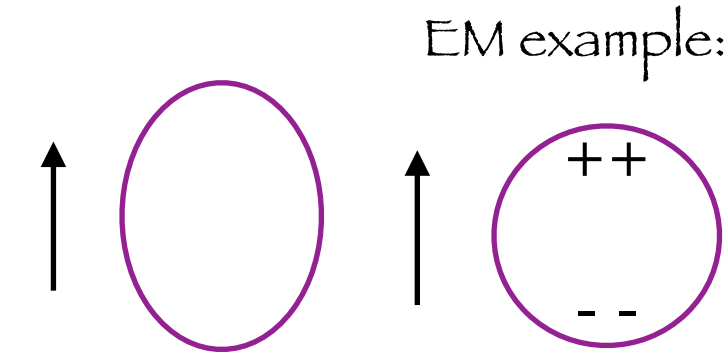
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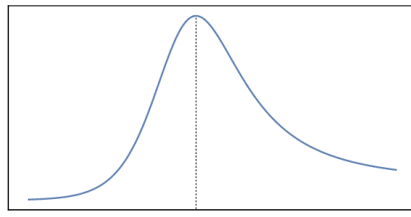


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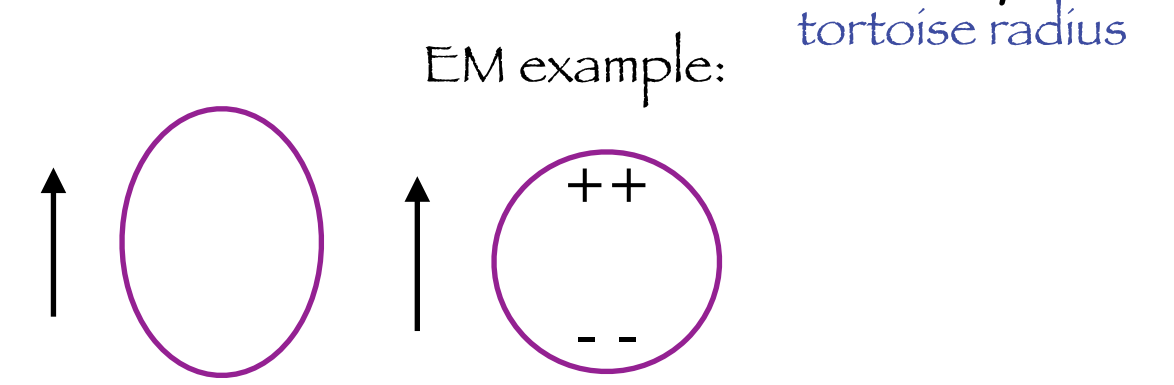
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the Love number surprise

Kerr metric:
$$ds^2 = -\frac{\rho^2 - r_s r}{\rho^2} dt^2 - \frac{2ar_s r \sin^2 \theta}{\rho^2} dt d\varphi + \frac{\rho^2}{\Delta} dr^2 + \rho^2 d\theta^2 + \frac{(r^2 + a^2)^2 - a^2 \Delta \sin^2 \theta}{\rho^2} \sin^2 \theta d\varphi^2$$

$$\Delta \equiv r^2 - rr_s + a^2 \quad \rho^2 \equiv r^2 + a^2 \cos^2 \theta$$

- Teukolsky equation:

$$\partial_r \left(\Delta \partial_r \phi_\ell^{(s)} \right) + s(2r - r_s) \partial_r \phi_\ell^{(s)} + \left(\frac{a^2 m^2 + is(2r - r_s)am}{\Delta} - (\ell - s)(\ell + s + 1) \right) \phi_\ell^{(s)} = 0$$

$$\Delta \equiv r^2 - rr_s + a^2 \quad \text{static: } \partial_t = 0 \quad \text{expand in spin-weighted spherical harmonics}$$

$$a = \text{BH spin} \quad , \quad r_s = 2GM \quad , \quad s = 0, 1, 2 \quad , \quad \ell, m = \text{ang. mom. quantum no.}$$

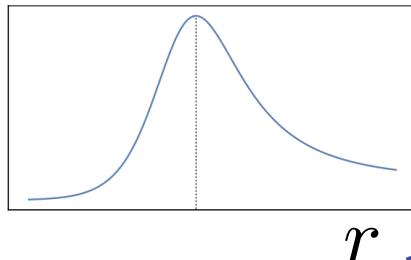
$\phi_\ell^{(s)}$ = Newman Penrose scalar \sim Weyl tensor projected onto null tetrad
 electric and magnetic types contained in real and imaginary parts thereof

Spin ladder operators connect spin 0, 1, 2 perturbations i.e. by taking suitable derivatives, can raise and lower spin. This can be viewed as a generalization of the Teukolsky-Starobinsky identities. For static perturbations, something special happens.

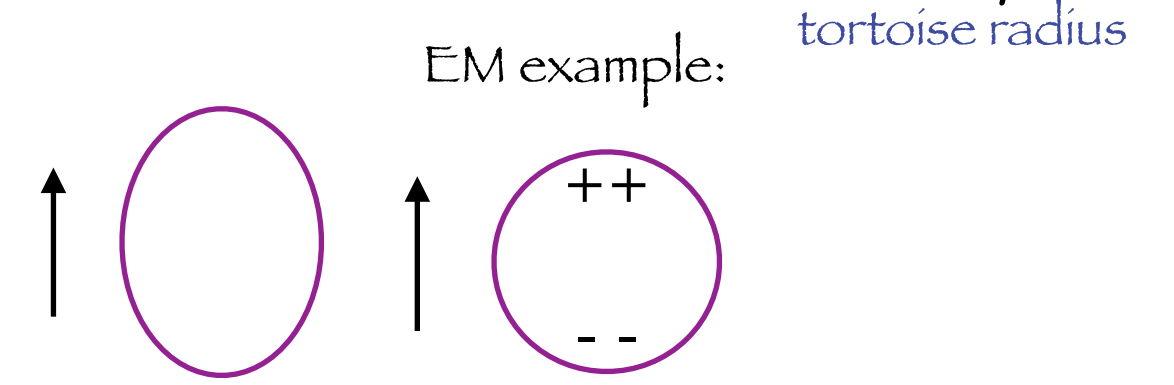
The upshot: it is sufficient to understand the spin 0 case: $\square \phi = 0$

And for pedagogy, let's focus on Schwarzschild (will comment on Kerr).

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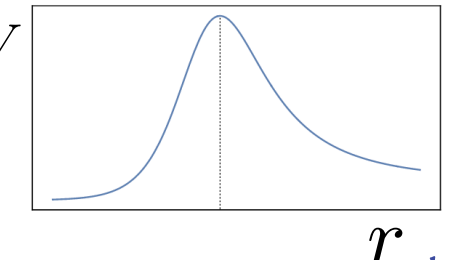
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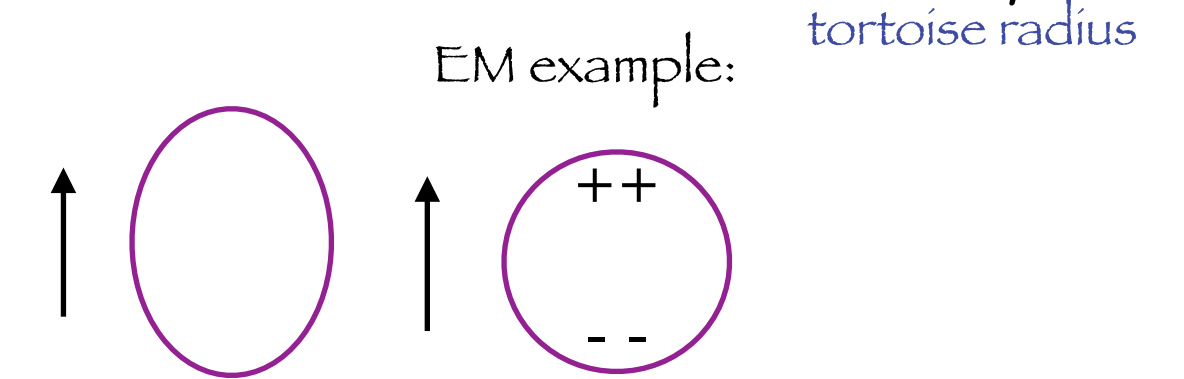
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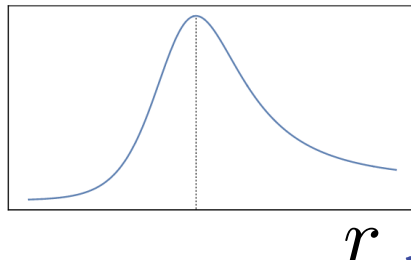
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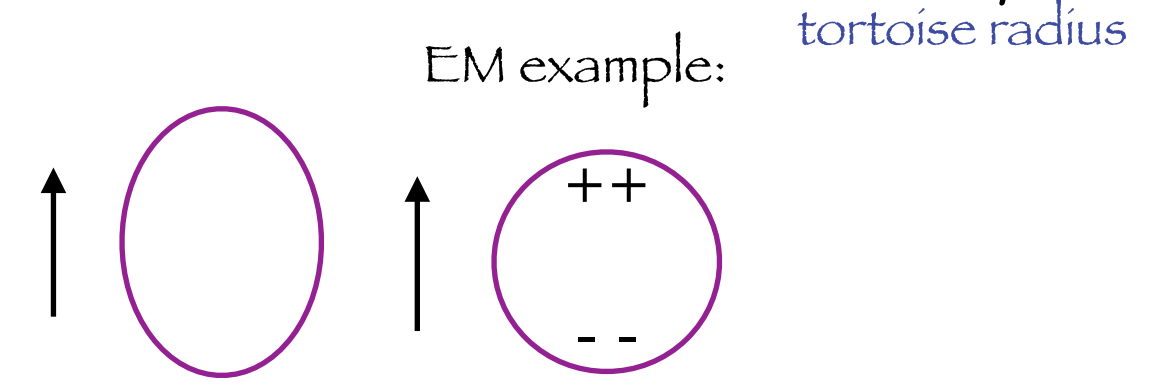
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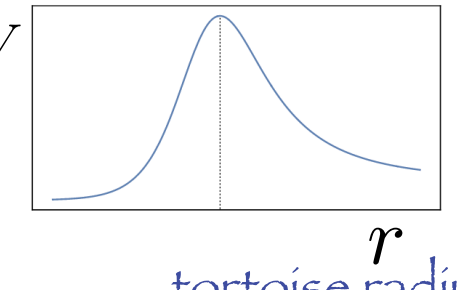
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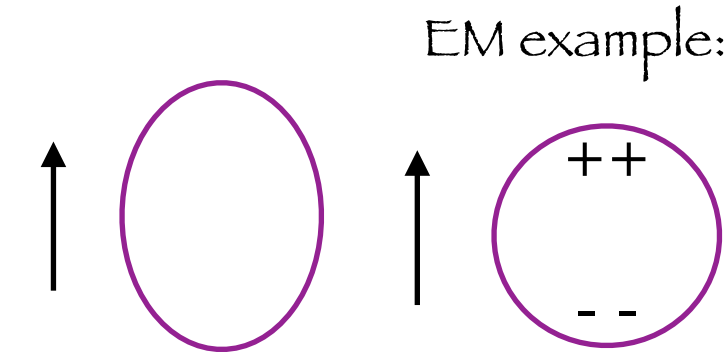
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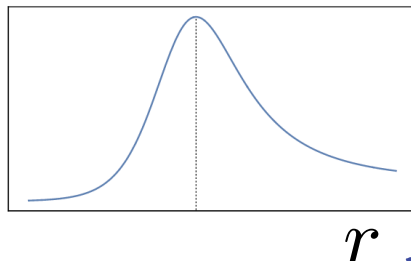
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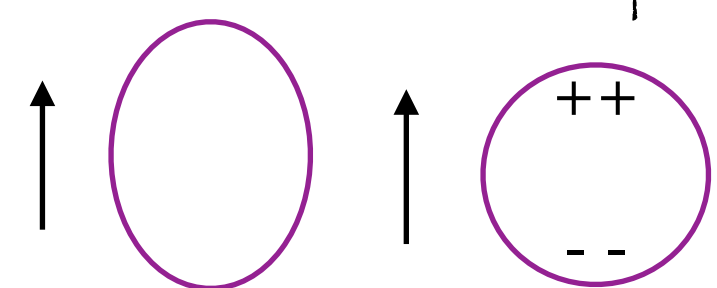
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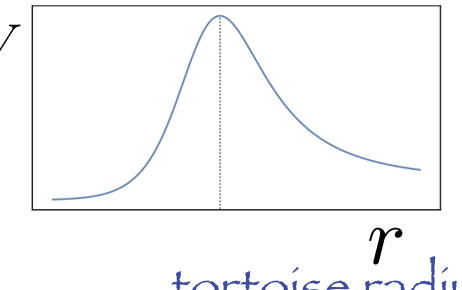
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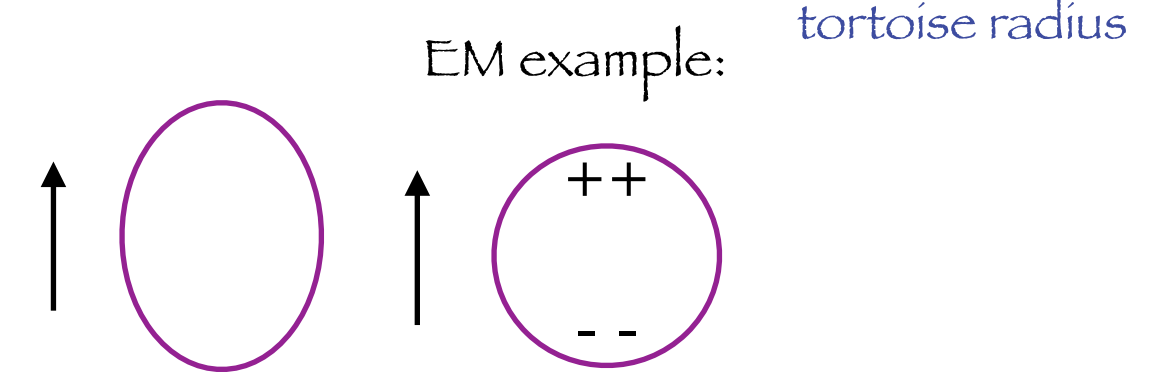
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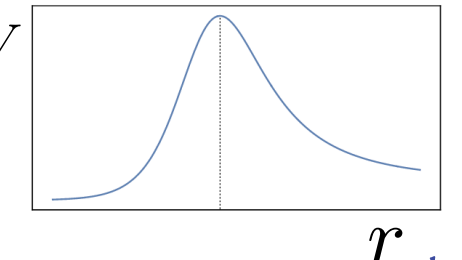
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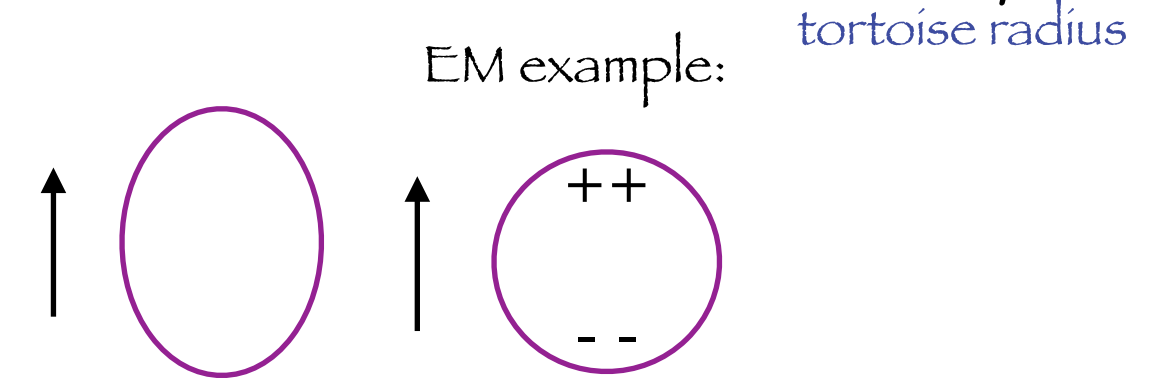
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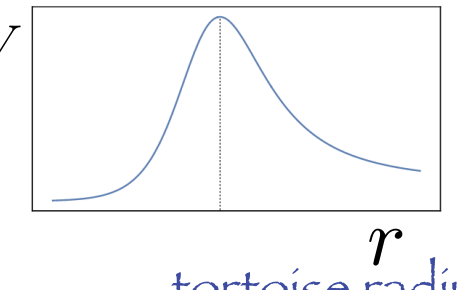
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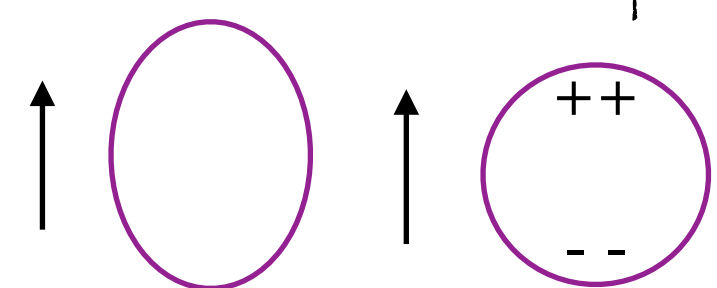
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- Extra comments: 1. $SO(3,1)$ symmetry is exact (for static perturbations). 2. At large r , \vec{K} generates special conformal trans. 3. Connection with reflection-less potential. 4. Higher symmetries: $\delta\phi_\ell = [D^+]^\ell \Delta \partial_r ([D^-]^\ell \phi_\ell)$.

Symmetries: summary and further thoughts

- For static scalar on Schwarzschild: $SO(3,1)$ exact symmetry. For Kerr: J_3, K_3 survive and are sufficient to generate the ladder structure for the Love number argument.
- For spin 1 and 2 perturbations: apply spin ladder.
- For low frequency perturbations, $SO(4,2)$ approximate symmetry is present, making possible the study of dynamical tidal response. An associated conserved charge is curiously connected with superradiance instability rate.
- For higher dimensional BH, ladder structure exists for $\hat{\ell} \equiv \frac{\ell}{D-3}$, and only for integer $\hat{\ell}$ can the “ground state” connection be made, which is crucial for the vanishing of the corresponding Love number.
- Nonlinear tidal response is being explored.

See also: Charalambous, Dubovsky, Ivanov; De Luca, Khoury, Wong; Riva, Santoní, Savíc, Vernízzí; Rodríguez, Santoní, Solomon, Temoche; Rai, Santoní

Nonlinearity: quadratic quasi-normal modes

- Linear perturbation theory: $g = g_{\text{BH}} + \phi \longrightarrow [\partial^2 + V]\phi \sim 0$
Second order perturbation theory: $[\partial^2 + V]\phi \sim \partial^2 \phi^2$

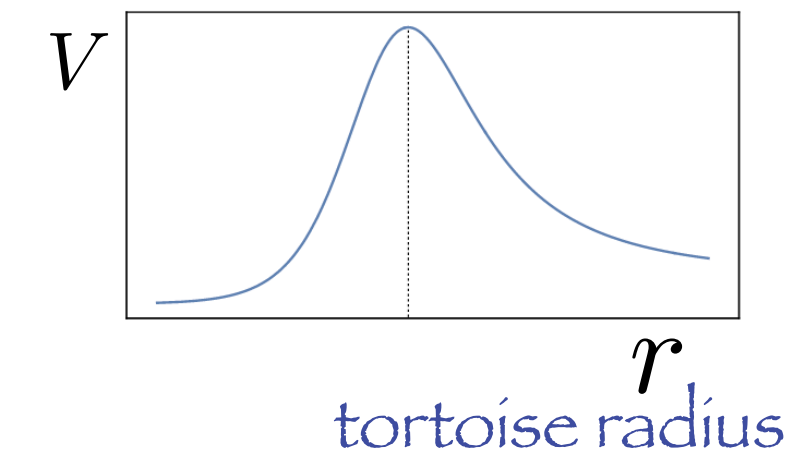
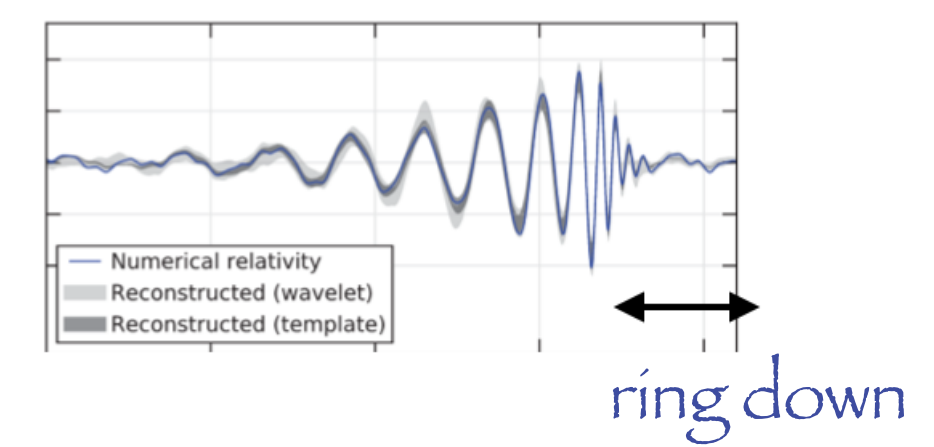
Write $\phi = \phi^{(1)} + \phi^{(2)} + \dots \longrightarrow$

$$[\partial^2 + V]\phi^{(1)} \sim 0$$
$$[\partial^2 + V]\phi^{(2)} \sim \partial^2 \phi^{(1)2}$$

Not surprising that: $\phi^{(1)}$ contains $e^{-i\omega_{220}t}$ \longrightarrow $\phi^{(2)}$ contains $e^{-i\underline{2}\omega_{220}t}$

linear QNM quadratic QNM

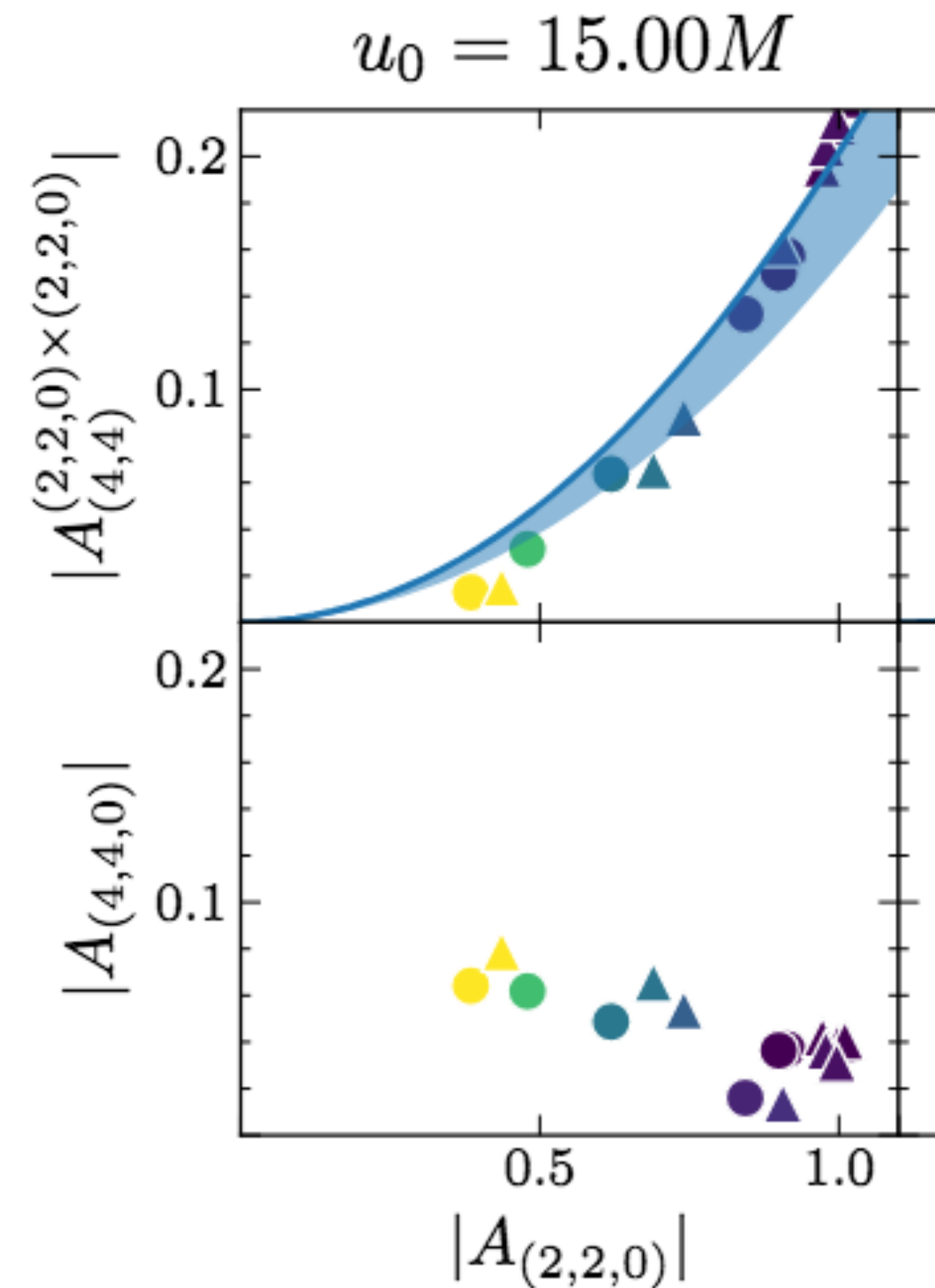
We searched for quadratic QNM in black hole merger simulations, and verified (220)x(220) amplitude goes as square of (220) amplitude (Mitman et al., Cheung et al. + much prior work).



Expect quadratic quasi-normal modes from pairs of linear quasi-normal modes.

e.g. from pairs of ω_{220} we get quadratic $\omega = 2\omega_{220}$

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Collaboration with Mitman, Lagos, Stein, Ma et al.

Collaboration with Lagos.

See also Cheung et al.

Nonlinearity: quadratic quasi-normal modes

- Linear perturbation theory: $g = g_{\text{BH}} + \phi \longrightarrow [\partial^2 + V]\phi \sim 0$
Second order perturbation theory: $[\partial^2 + V]\phi \sim \partial^2 \phi^2$

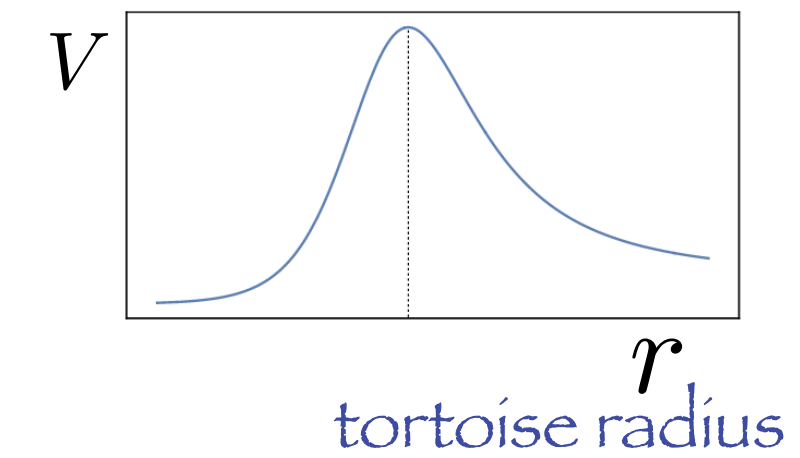
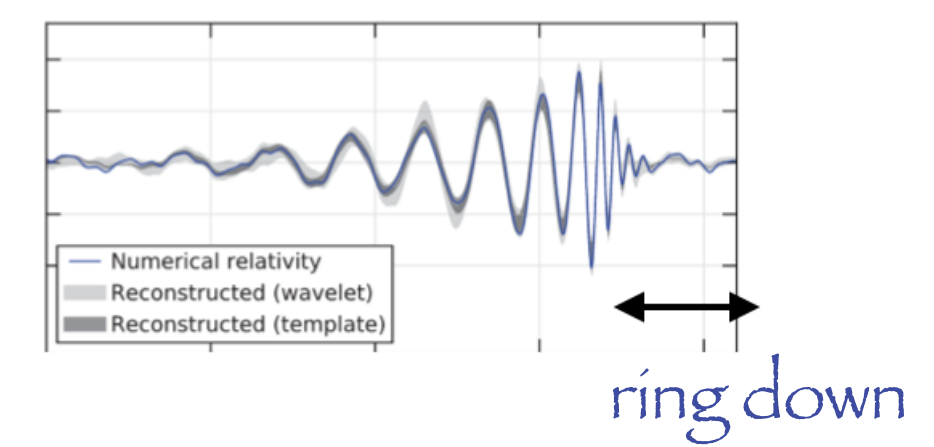
Write $\phi = \phi^{(1)} + \phi^{(2)} + \dots \longrightarrow$

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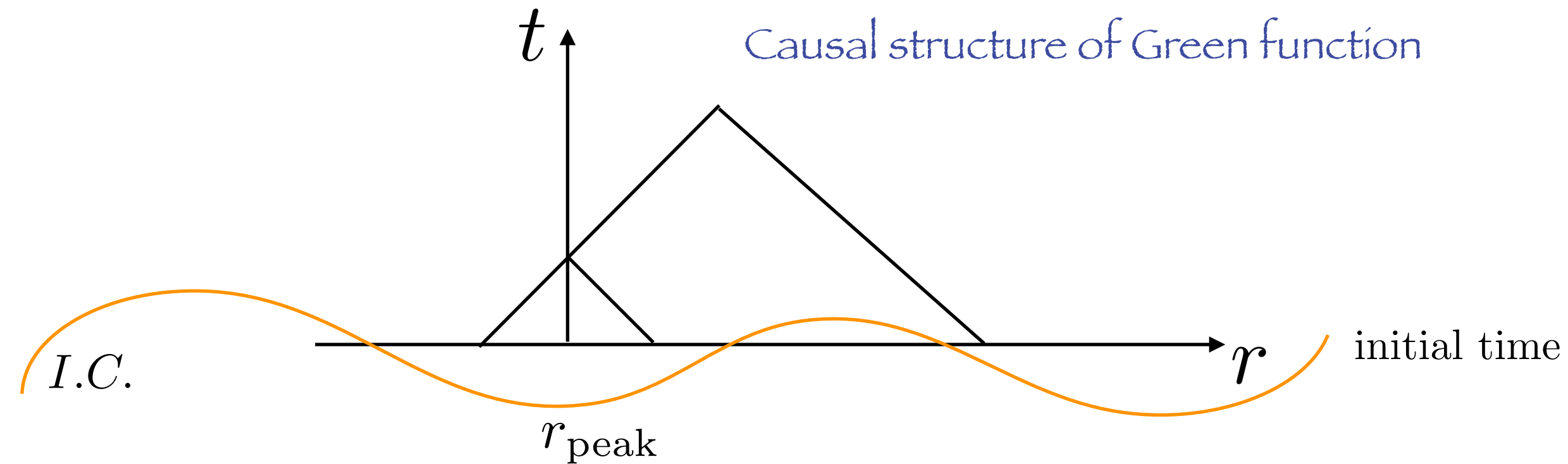
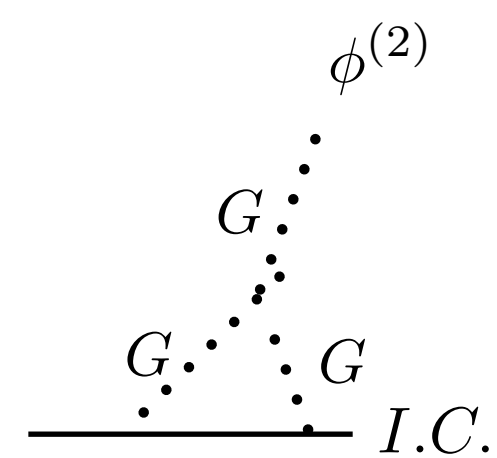
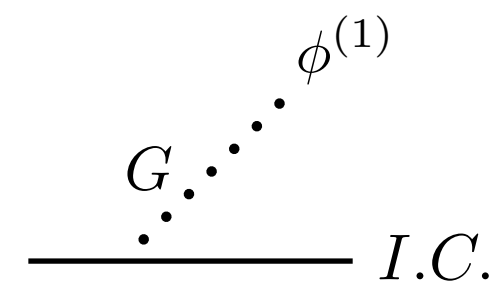
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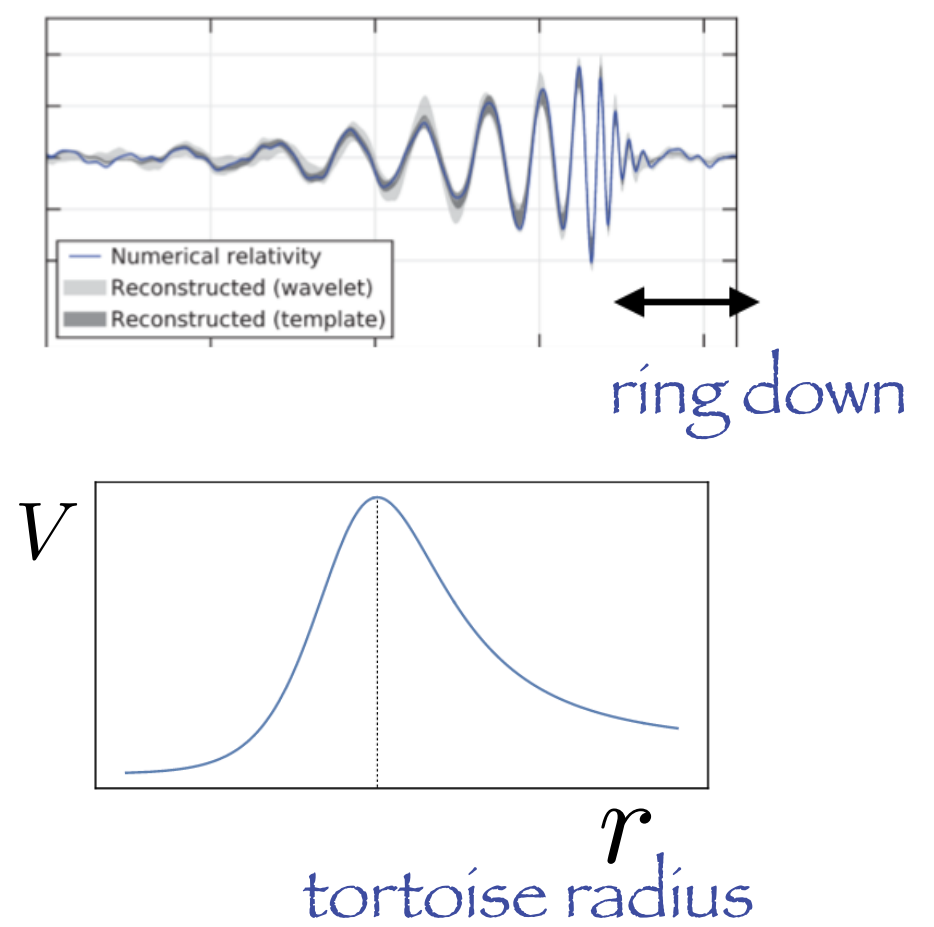
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Many (known) implications: e.g. (1) ring-down waveform is not purely superposition of QNMs, (2) amplitude of QNMs is time dependent, (3) QNMs are generated from I.C. by scattering off potential barrier, etc.



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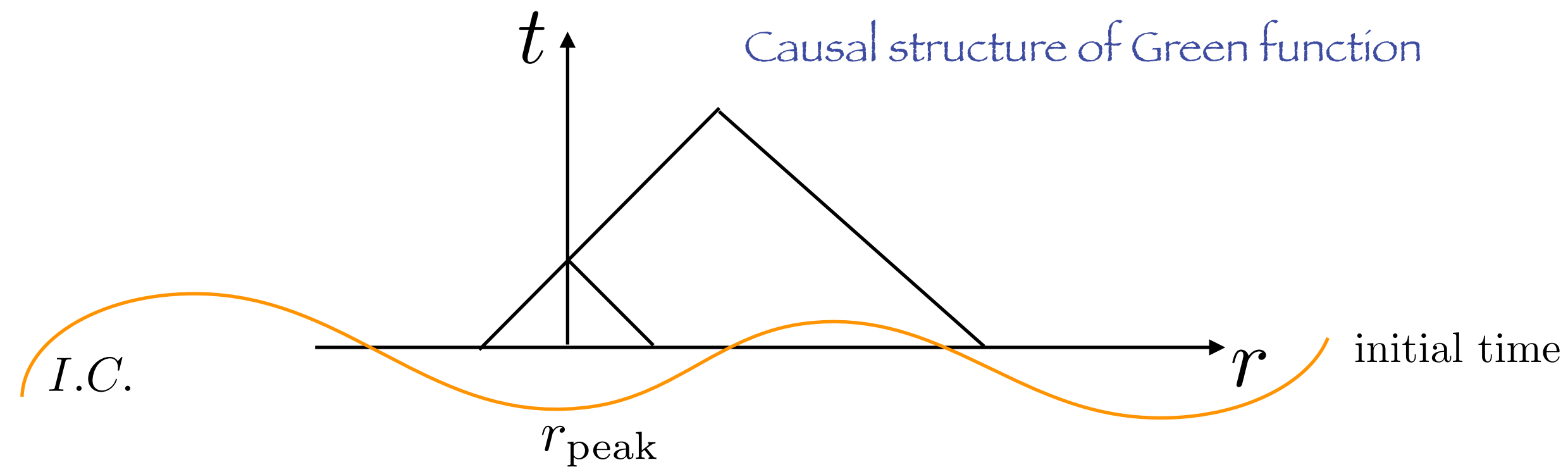
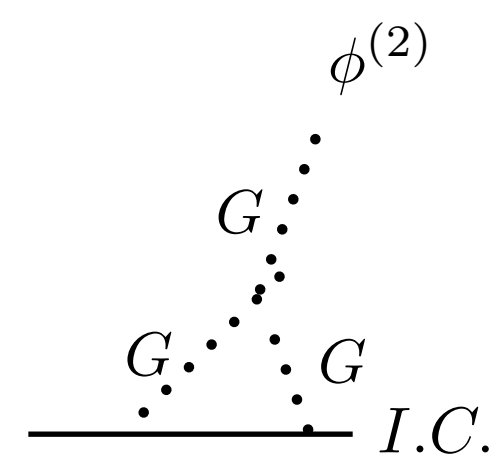
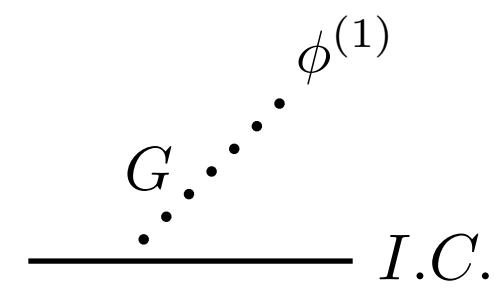
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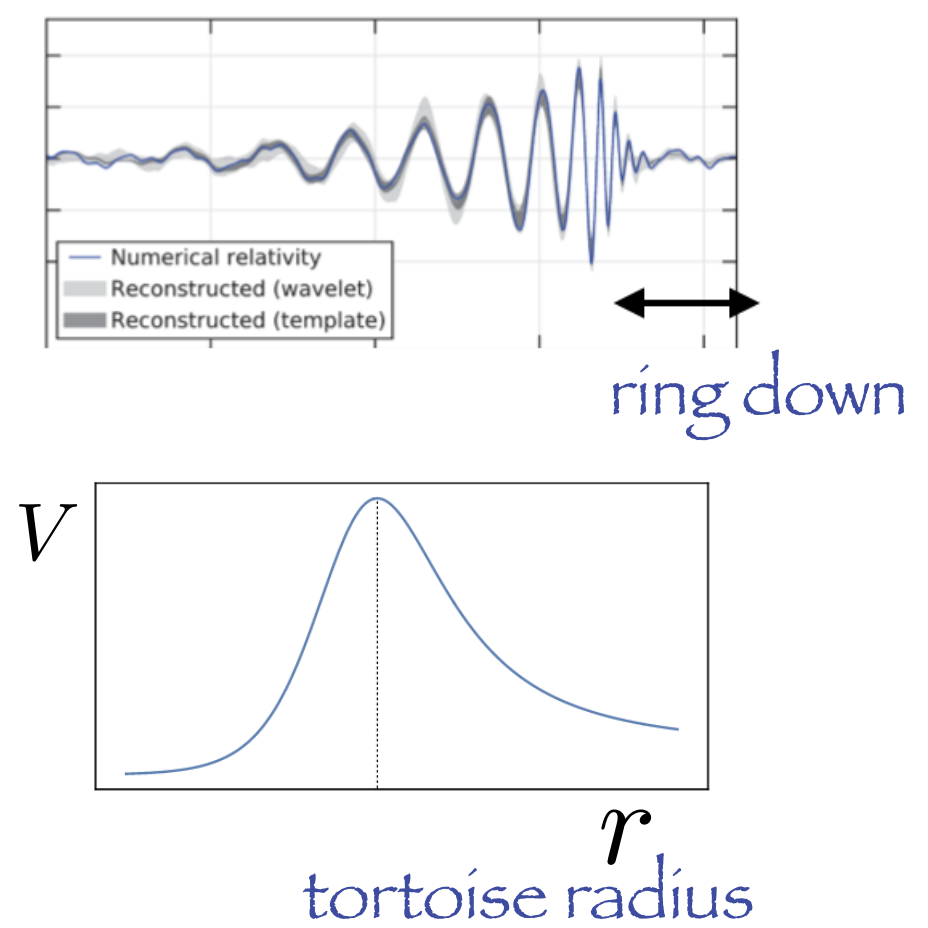
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- Why should we care? (1) more accurate modeling of the ring-down, (2) may get us closer to merger time, (3) relation between $\phi^{(2)}$ and $\phi^{(1)}$ (including amplitude) provides test of GR (e.g. Kehagias/Riotto, Redondo-Yuste+, Zhu+)

Ladder structure from $SO(3,1)$ explains vanishing Love number

Quadratic QNMs are detectable and interesting



*Symmetry and nonlinearity
in black hole perturbation theory*

Lam Hui

Columbia University

Work done with: Joyce, Penco, Santoni, Solomon; Beren, Sun; Lagos; Mittman, Lagos, Stein+; Podo